Candidate surname	elow before ento	Other names
Pearson Edexcel International Advanced Level	entre Number	Candidate Number
Wednesday 22	May	2019
Afternoon (Time: 1 hour 20 minutes)) Paper R	Reference WPH06/01
Physics Advanced Unit 6: Experimental Phy	sics	
	: Tota	

Instructions

- Use **black** ink or ball-point pen.
- Fill in the boxes at the top of this page with your name, centre number and candidate number.
- Answer all questions.
- Answer the questions in the spaces provided
 - there may be more space than you need.

Information

- The total mark for this paper is 40.
- The marks for **each** question are shown in brackets
 - use this as a guide as to how much time to spend on each question.
- The list of data, formulae and relationships is printed at the end of this booklet.
- Candidates may use a scientific calculator.

Advice

- Read each question carefully before you start to answer it.
- Try to answer every question.
- Check your answers if you have time at the end.

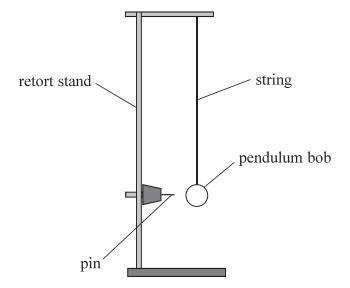
Turn over ▶





Answer ALL questions. Write your answers in the spaces provided.

A student investigated how the time period T of a simple pendulum varies with the length l of the pendulum, using the apparatus shown.



The pendulum performs simple harmonic motion, hence T is given by the equation

$$T = 2\pi \sqrt{\frac{l}{g}}$$

The student displaced the pendulum bob by a small distance to set the pendulum oscillating. The student measured the time for 10 complete oscillations.

(a) (i) State the reason for using a small displacement.

(1)

(ii) The student used a pin as a timing marker.

Explain how placing the pin at the centre of oscillation would lead to a more accurate value for 10T.

(2)

(b) The student obtained the following results.

10 <i>T</i> /s	mean $10T/s$	mean T/s				
10.15	10.32	10.20	10.28	10.24		

(i) Complete the table.

(2)

(ii) Calculate the percentage uncertainty in T.

(2)

Percentage uncertainty =

(c) The student doubled the length of the pendulum. She made the following prediction.



Prediction:

If time period for short pendulum = T_s and time period for double length = T_d Then T_d^2 / T_s^2 should be 2

(i) Justify the prediction.

(2)



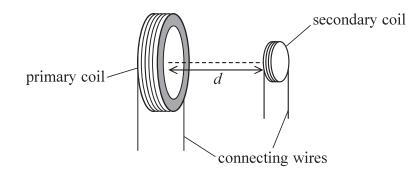
(ii) The student determines T for the longer pendulum as 1.461 \pm Determine whether her results support her prediction.	0.011 s. (4)
(Total for Q	uestion 1 = 13 marks)



- An alternating current in a primary coil generates a magnetic field. This induces an e.m.f. across a small secondary coil placed in the magnetic field.
 - (a) Explain how an e.m.f. is induced across the secondary coil.

(2)

(b) A student has the apparatus shown.



The e.m.f. V induced across the secondary coil is given by

$$V = \frac{kNI}{d^3}$$

where

k is a constant,

N is the number of turns on the primary coil,

I is the current in the primary coil,

d is the distance between the centres of the coils as shown.



The student investigates how V varies with d .	
Write a plan to determine k using a graphical method.	
Your plan should include:	
(i) a circuit diagram to show how the primary coil should be connected with any additional components required,	
(1)
(ii) the measurements to be made with any additional apparatus required, (3)
(iii) the graph to be plotted and how it would be used to determine k , (2))
(iv) a statement of a significant source of uncertainty. (1)



(T	Total for Question 2 = 9 marks)



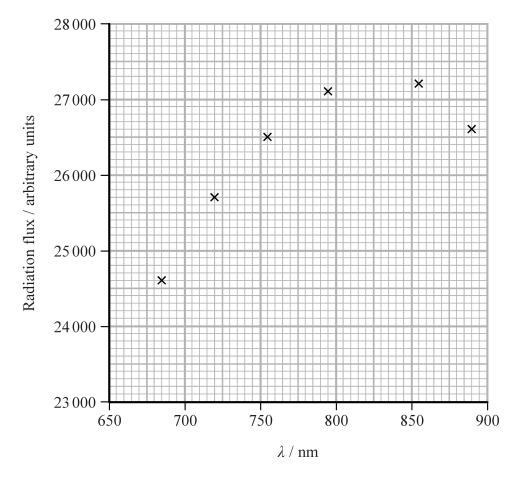
3 The surface temperature T of a star can be estimated by determining the wavelength λ_{\max} at which peak power emission occurs from the star.

A student astronomer observed the red giant Betelgeuse.

He measured the radiation flux at six wavelengths and obtained the following results.

Radiation flux / arbitrary units	λ / nm
24600	685
25 700	720
26500	755
27 100	795
27200	855
26600	890

The data was plotted as shown.

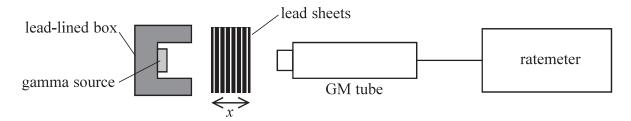


(a) (i) Draw a line of best fit to show where the radiation flux is a maximum. **(2)** (ii) Estimate λ_{max} . (1) (iii) Calculate a value for T. (2) (b) Suggest two reasons why the value of λ_{\max} might not be accurate. **(2)**

(Total for Question 3 = 7 marks)



4 A student investigated the absorption of gamma radiation by lead, using the apparatus shown.



(a) She recorded and corrected for the background count rate.

Explain why it is possible to correct for background count rate.

(2)

(b) The student recorded the corrected count rate C with different thicknesses of lead sheets.

The relationship between C and the total thickness x of the lead sheets is

$$C = C_0 e^{-\mu x}$$

where C_0 is the count rate without any lead sheets and μ is a constant.

Explain why plotting a graph of ln C against x will produce a straight line.

(2)

(c) The student obtained the following results.

<i>x</i> / mm	C/s^{-1}
1.52	132
3.89	112
6.81	95
9.33	86
11.48	74
13.70	67

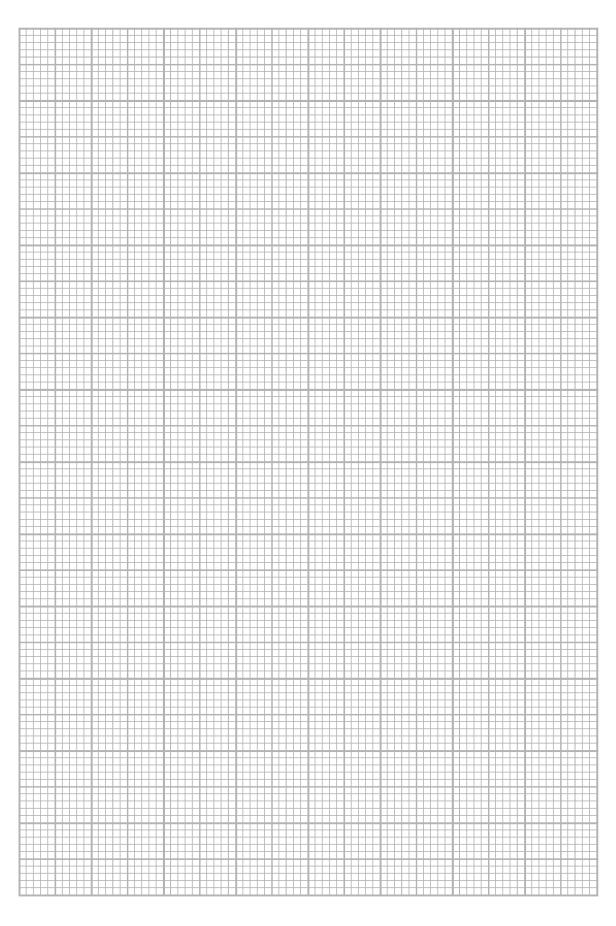
(i) Plot a graph of $\ln C$ against x on the grid provided. Use the additional column in the results table to record your processed data.

(5)

(ii) Determine a value for μ .

(2)

 $\mu =$



(Total for Question 4 = 11 marks)

TOTAL FOR PAPER = 40 MARKS





List of data, formulae and relationships

Acceleration of free fall
$$g = 9.81 \text{ m s}^{-2}$$
 (close to Earth's surface)

Boltzmann constant
$$k = 1.38 \times 10^{-23} \text{ J K}^{-1}$$

Coulomb's law constant
$$k = 1/4\pi\varepsilon_0$$

$$= 8.99 \times 10^9 \text{ N m}^2 \text{ C}^{-2}$$

Electron charge
$$e = -1.60 \times 10^{-19} \,\mathrm{C}$$

Electron mass
$$m_{e} = 9.11 \times 10^{-31} \,\mathrm{kg}$$

Electronvolt
$$1 \text{ eV} = 1.60 \times 10^{-19} \text{ J}$$

Gravitational constant
$$G = 6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$$

Gravitational field strength
$$g = 9.81 \text{ N kg}^{-1}$$
 (close to Earth's surface)

Permittivity of free space
$$\varepsilon_0 = 8.85 \times 10^{-12} \, \mathrm{F m}^{-1}$$

Planck constant
$$h = 6.63 \times 10^{-34} \,\mathrm{J s}$$

Proton mass
$$m_{\rm p} = 1.67 \times 10^{-27} \text{ kg}$$

Speed of light in a vacuum
$$c = 3.00 \times 10^8 \,\mathrm{m \, s^{-1}}$$

Stefan-Boltzmann constant
$$\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$

Unified atomic mass unit
$$u = 1.66 \times 10^{-27} \text{ kg}$$

Unit 1

Mechanics

Kinematic equations of motion
$$v = u + at$$

$$s = ut + \frac{1}{2}at^2$$
$$v^2 = u^2 + 2as$$

Forces $\Sigma F = ma$

$$g = F/m$$
$$W = mg$$

Work and energy $\Delta W = F \Delta s$

$$E_{\rm k} = \frac{1}{2}mv^2$$

$$\Delta E_{\rm grav} = mg\Delta h$$

Materials

Stokes' law $F = 6\pi \eta r v$

Hooke's law $F = k\Delta x$

Density $\rho = m/V$

Pressure p = F/A

Young modulus $E = \sigma/\varepsilon$ where

Stress $\sigma = F/A$ Strain $\varepsilon = \Delta x/x$

Elastic strain energy $E_{el} = \frac{1}{2}F\Delta x$



Unit 2

Waves

Wave speed $v = f\lambda$

Refractive index $\mu_2 = \sin i / \sin r = v_1 / v_2$

Electricity

Potential difference V = W/Q

Resistance R = V/I

Electrical power, energy and P = VI efficiency $P = I^2 K$

 $P = I^2 R$ $P = V^2 / R$

W = VIt

% efficiency =
$$\frac{\text{useful energy output}}{\text{total energy input}} \times 100$$

% efficiency =
$$\frac{\text{useful power output}}{\text{total power input}} \times 100$$

Resistivity $R = \rho l/A$

Current $I = \Delta Q/\Delta t$

I = nqvA

Resistors in series $R = R_1 + R_2 + R_3$

Resistors in parallel $\frac{1}{R} = \frac{1}{R_1} + \frac{1}{R_2} + \frac{1}{R_3}$

Quantum physics

Photon model E = hf

Einstein's photoelectric $hf = \phi + \frac{1}{2}mv_{\text{max}}^2$

equation



Unit 4

Mechanics

Momentum p = mv

Kinetic energy of a

non-relativistic particle $E_k = p^2/2m$

Motion in a circle $v = \omega r$

 $T=2\pi/\omega$

 $F = ma = mv^2/r$

 $a = v^2/r$

 $a = r\omega^2$

Fields

Coulomb's law $F = kQ_1Q_2/r^2$ where $k = 1/4\pi\epsilon_0$

Electric field E = F/Q

 $E = kQ/r^2$

E = V/d

Capacitance C = Q/V

Energy stored in capacitor $W = \frac{1}{2}QV$

Capacitor discharge $Q = Q_0 e^{-t/RC}$

In a magnetic field $F = BII \sin \theta$

 $F = Bqv \sin \theta$

r = p/BQ

Faraday's and Lenz's laws $\varepsilon = -d(N\phi)/dt$

Particle physics

Mass-energy $\Delta E = c^2 \Delta m$

de Broglie wavelength $\lambda = h/p$



Unit 5

Energy and matter

Heating $\Delta E = mc\Delta\theta$

Molecular kinetic theory $\frac{1}{2}m\langle c^2 \rangle = \frac{3}{2}kT$

Ideal gas equation pV = NkT

Nuclear Physics

Radioactive decay $dN/dt = -\lambda N$

 $\lambda = \ln 2/t_{1/2}$

 $N = N_0 e^{-\lambda t}$

Mechanics

Simple harmonic motion $a = -\omega^2 x$

 $a = -A\omega^2 \cos \omega t$ $v = -A\omega \sin \omega t$ $x = A\cos \omega t$ $T = 1/f = 2\pi/\omega$

Gravitational force $F = Gm_1m_2/r^2$

Observing the universe

Radiant energy flux $F = L/4\pi d^2$

Stefan-Boltzmann law $L = \sigma T^4 A$

 $L=4\pi r^2\sigma T^4$

Wien's law $\lambda_{\max} T = 2.898 \times 10^{-3} \text{ m K}$

Redshift of electromagnetic

radiation $z = \Delta \lambda / \lambda \approx \Delta f / f \approx v / c$

Cosmological expansion $v = H_0 d$